

Spring 2020 HamSCI Workshop

Mar 20-21, 2020

The University of Scranton

Scranton, PA

Using Amateur Radio to Validate Model-Based Properties of Earth's Protective Shield

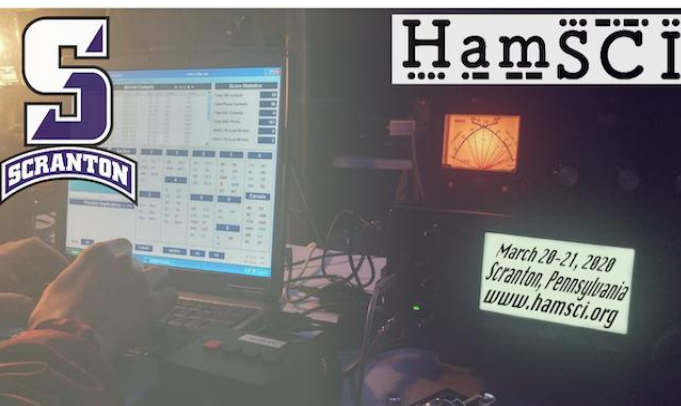
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Utah State University, Logan, UT

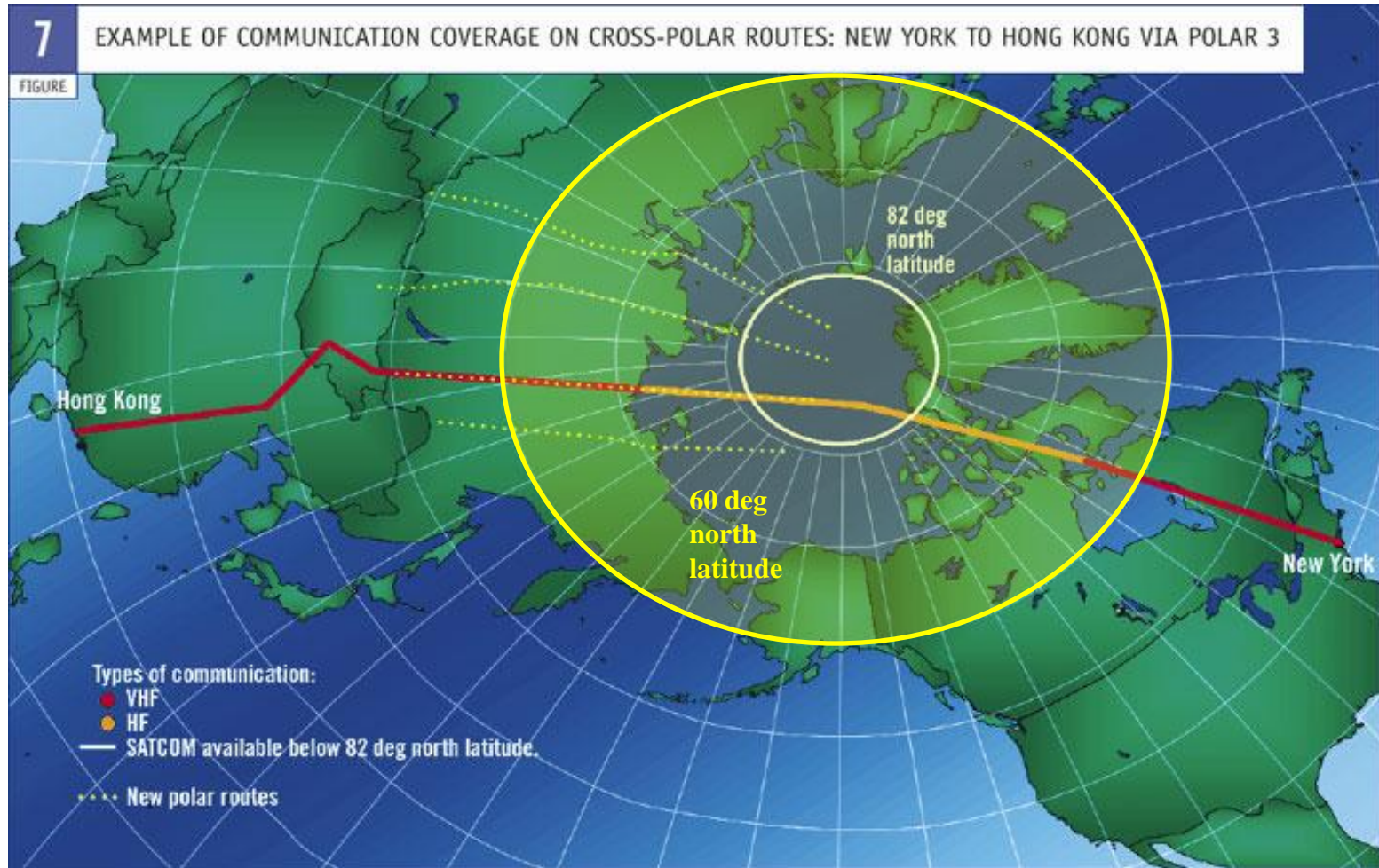


Introduction

Current Research Focus

- Polar Cap Absorption Events
 - Energetic Proton Transport
 - D-Region
 - HF Absorption

Introduction



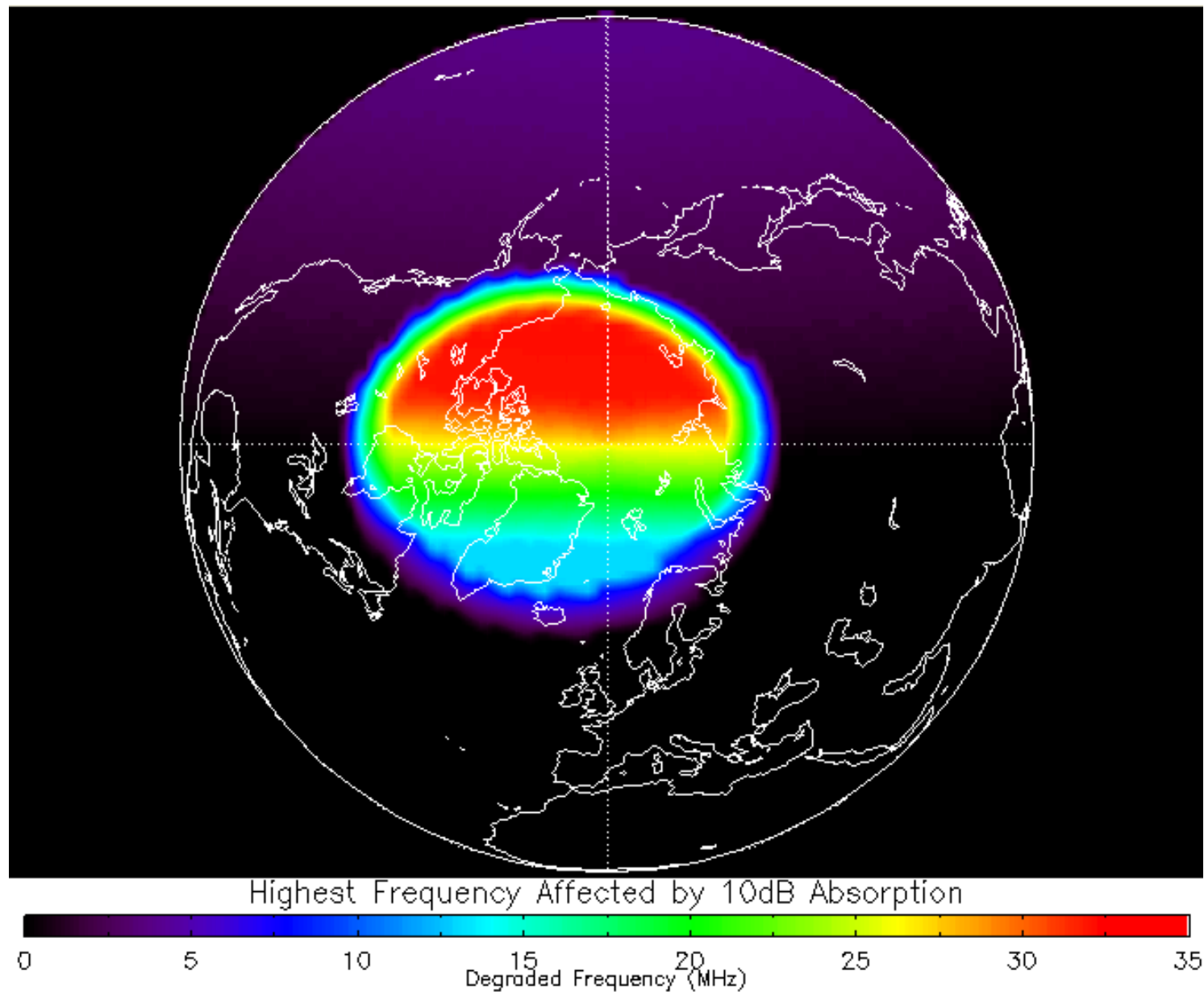
Average cost for diverting a flight between North America and Asia about \$100,000

Fiori & Danskin (2016)

As of 2007 about 7000 transpolar flights per year

Sauer & Wilkinson (2008)

Introduction



Elevated X-ray flux

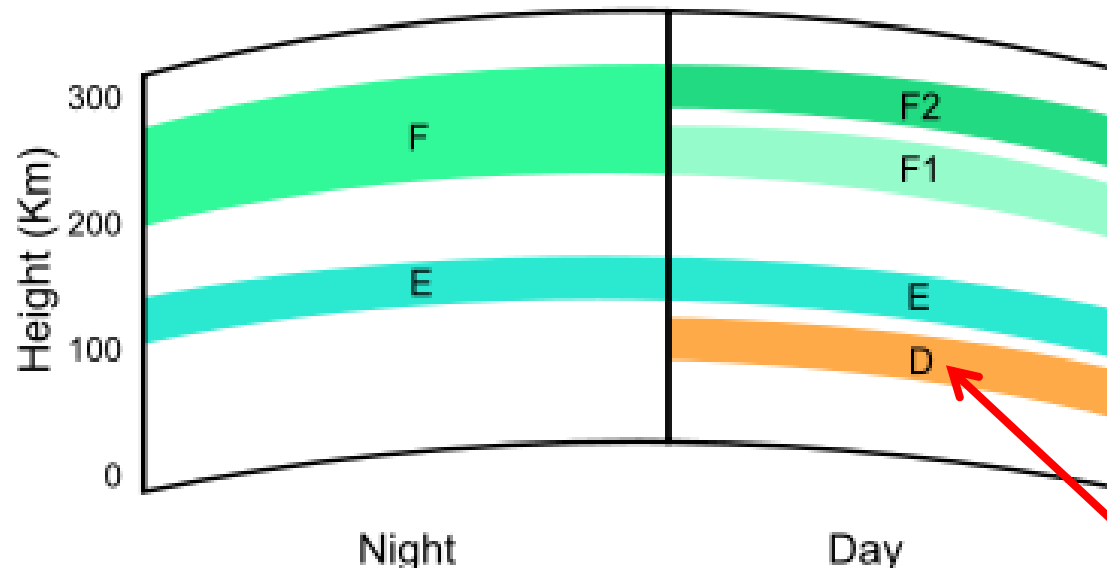
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Moderate Proton Flux

NOAA/SWPC Boulder, CO USA

Polar Cap Absorption Events (PCA)

The Ionosphere

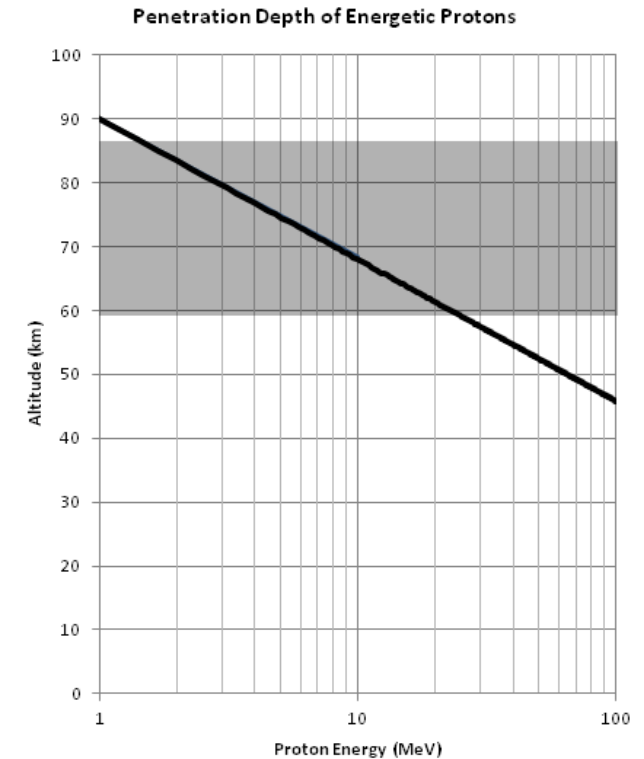
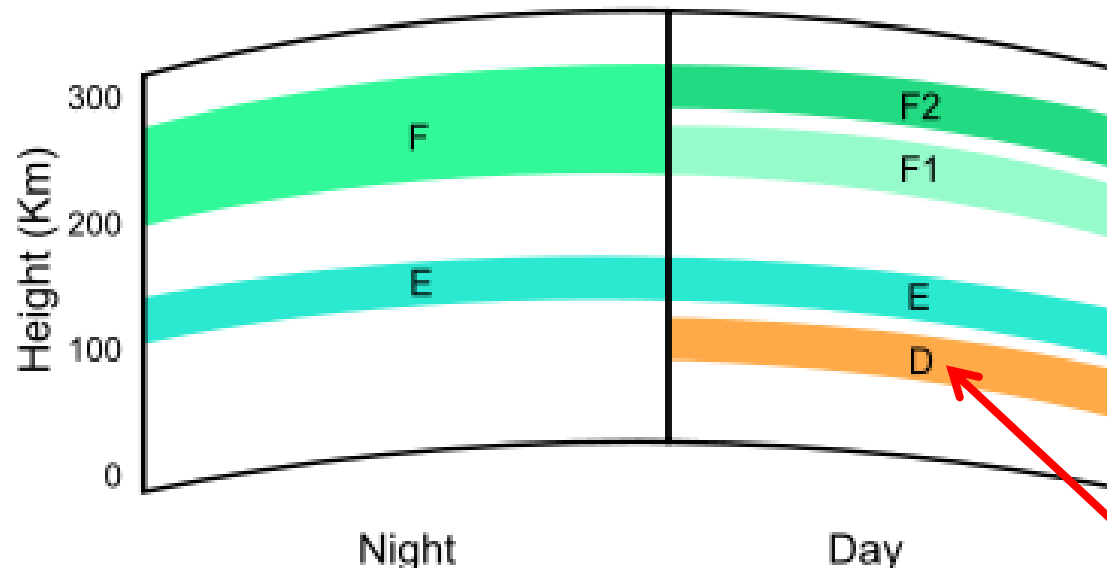


The D-Region

- *Altitude about 50-100 km*
- *Primarily Absorbs HF radio waves (3-30 MHz)*
- *Usually exists only during daylight*

Polar Cap Absorption Events (PCA)

The Ionosphere

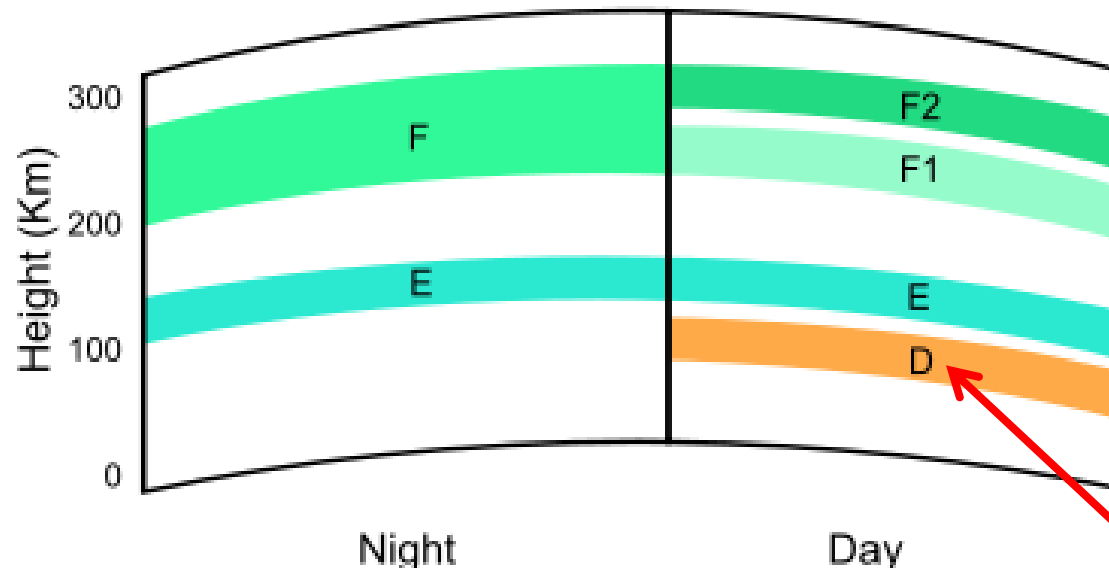


The D-Region

- Altitude about 50-100 km
- Primarily Absorbs HF radio waves (3-30 MHz)
- Usually exists only during daylight
- ***Enhanced ionization from solar protons during PCA events***
- ***1-20 MeV protons most responsible for PCA events***

Polar Cap Absorption Events (PCA)

The Ionosphere



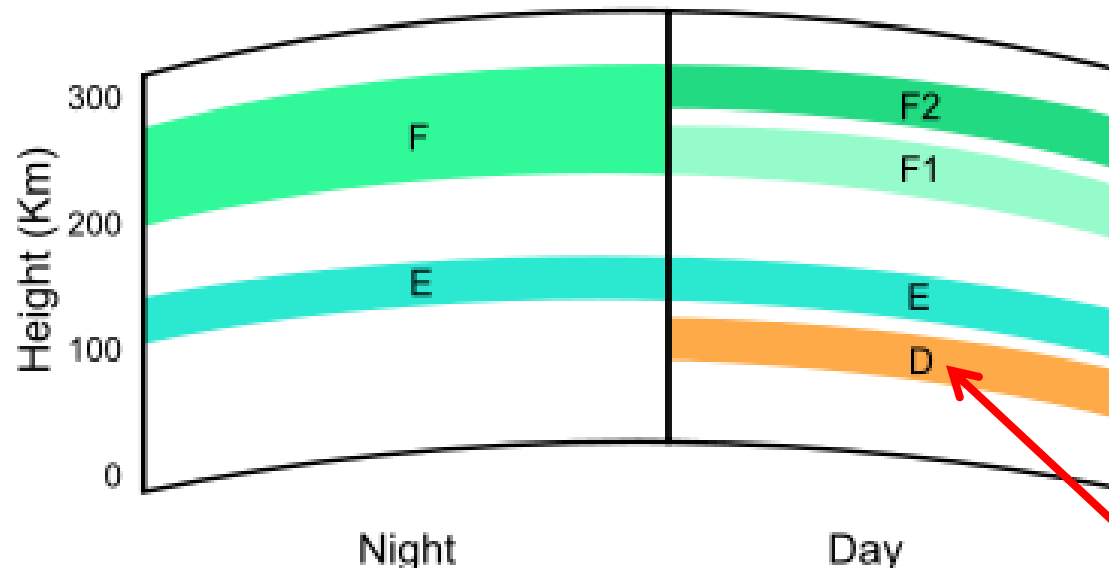
- Can solar protons reach any location?

The D-Region

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Polar Cap Absorption Events (PCA)

The Ionosphere



- Can solar protons reach any location?
- How do we study solar protons?

The D-Region

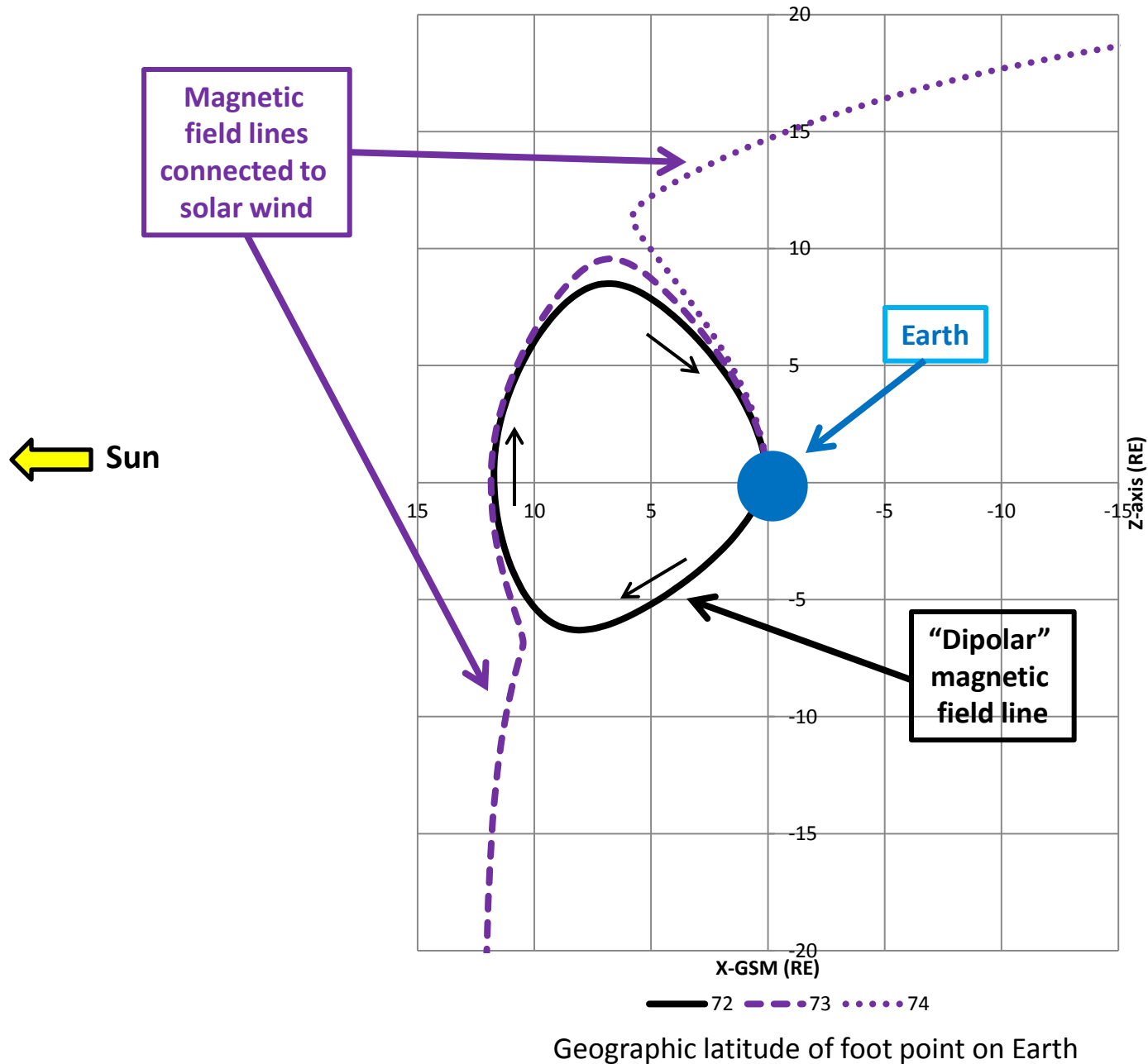
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Our Tracing Model

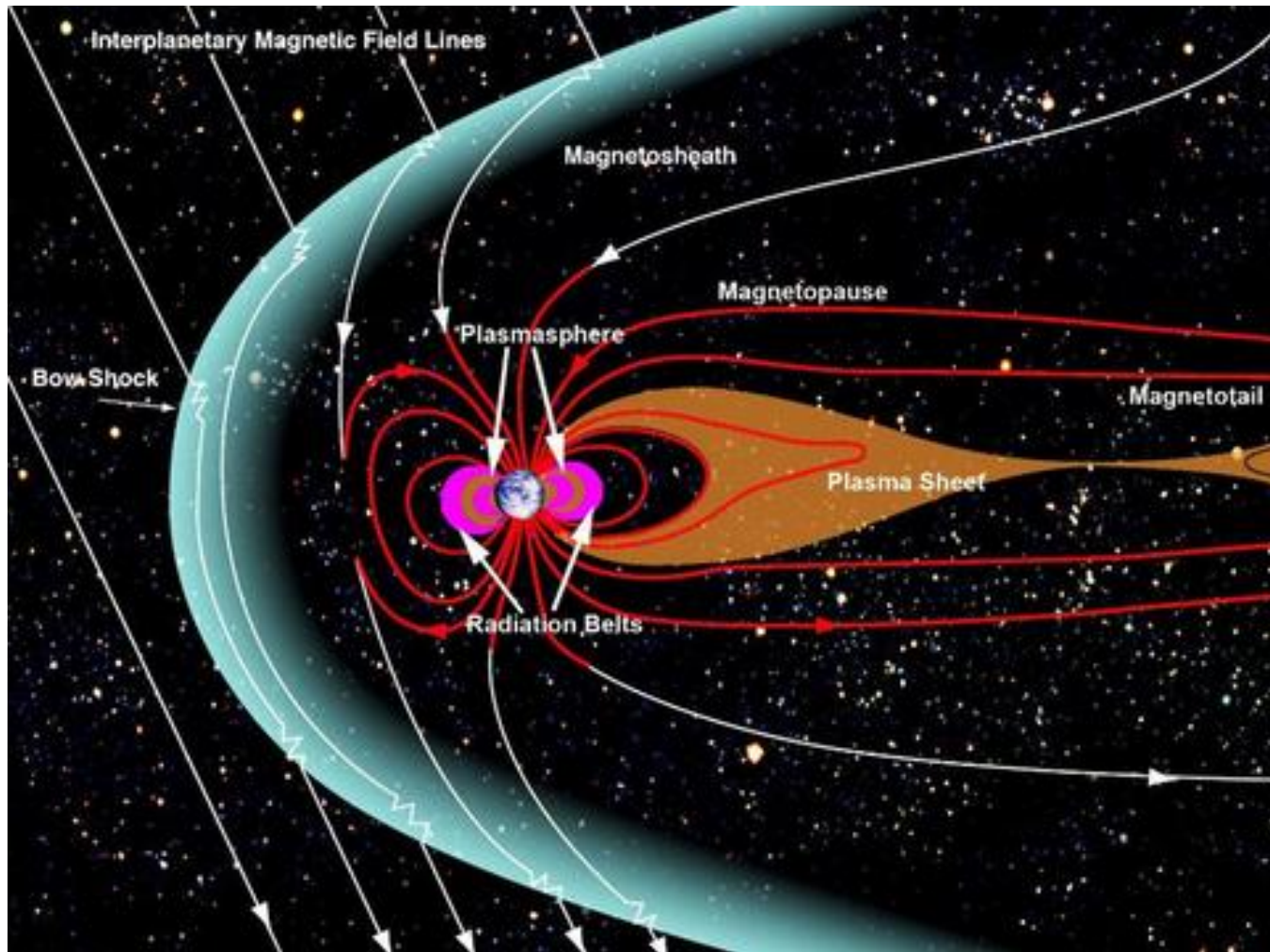
Tsyganenko Geomagnetic Field Model

- Semi-empirical best-fit representations of geomagnetic field
- Based on large number of satellite observations
- Outputs include:
 - Field line tracing
 - Geomagnetic field vector at selected point
- Internal Field – International Geomagnetic Reference Field (IGRF)
- External Field – 1996 Version of Tsyganenko Model (T96)

How Do We Study Solar Protons?



How Do We Study Solar Protons?



Open field lines connect Earth's upper atmosphere with Solar Wind

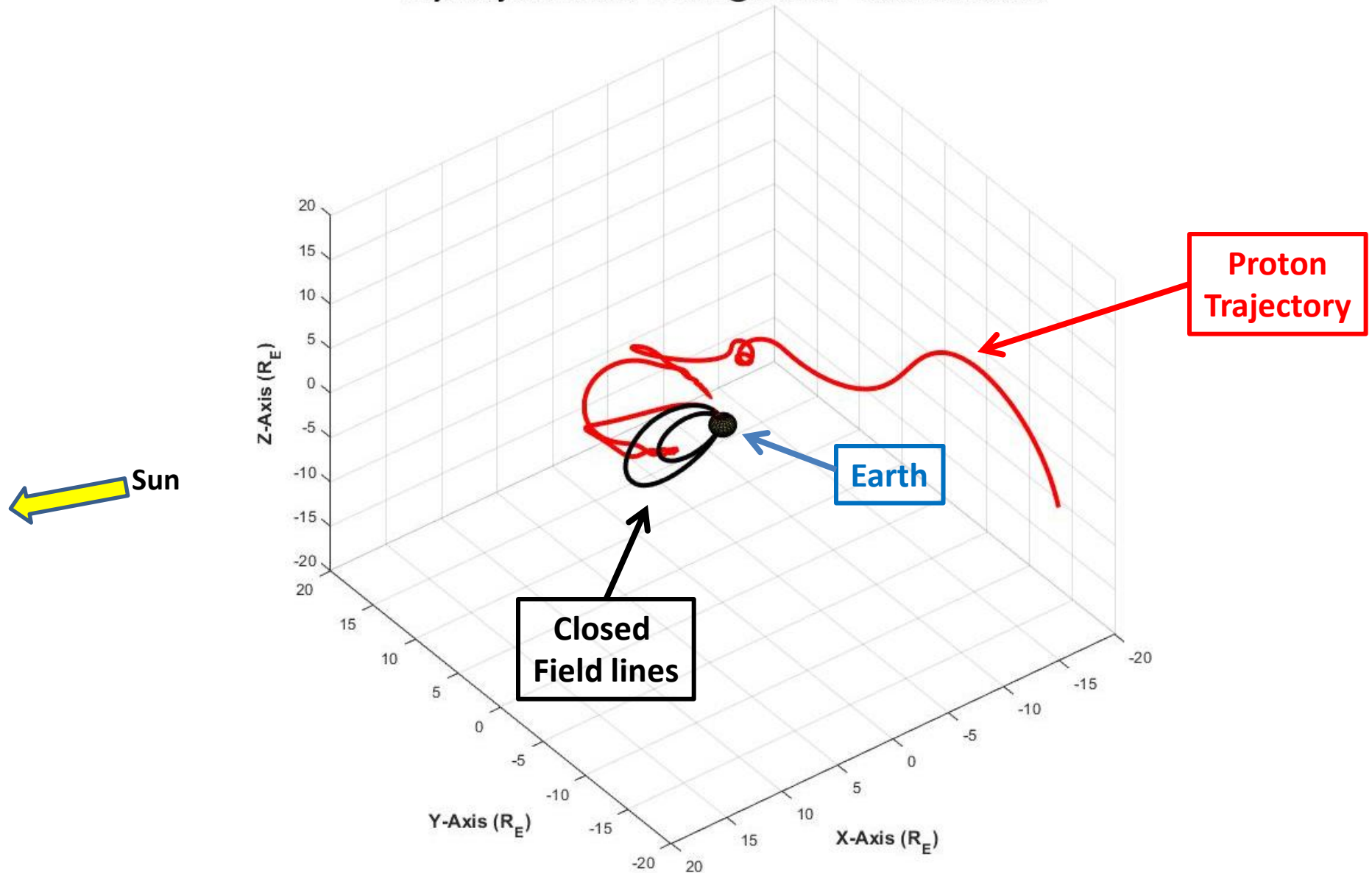
Can Solar Protons Reach Any Location?

Proton Energy Cutoff Latitude

- *Lowest latitude to which a solar proton can penetrate*
- Variable depending on
 - *Proton Energy*
 - *Local Time*
 - *Geomagnetic storm conditions*
 - *Universal Time*

Proton Energy Cutoff Latitude

Trajectory of 17.8 MeV Proton @ 0500 UT - Quiet Conditions



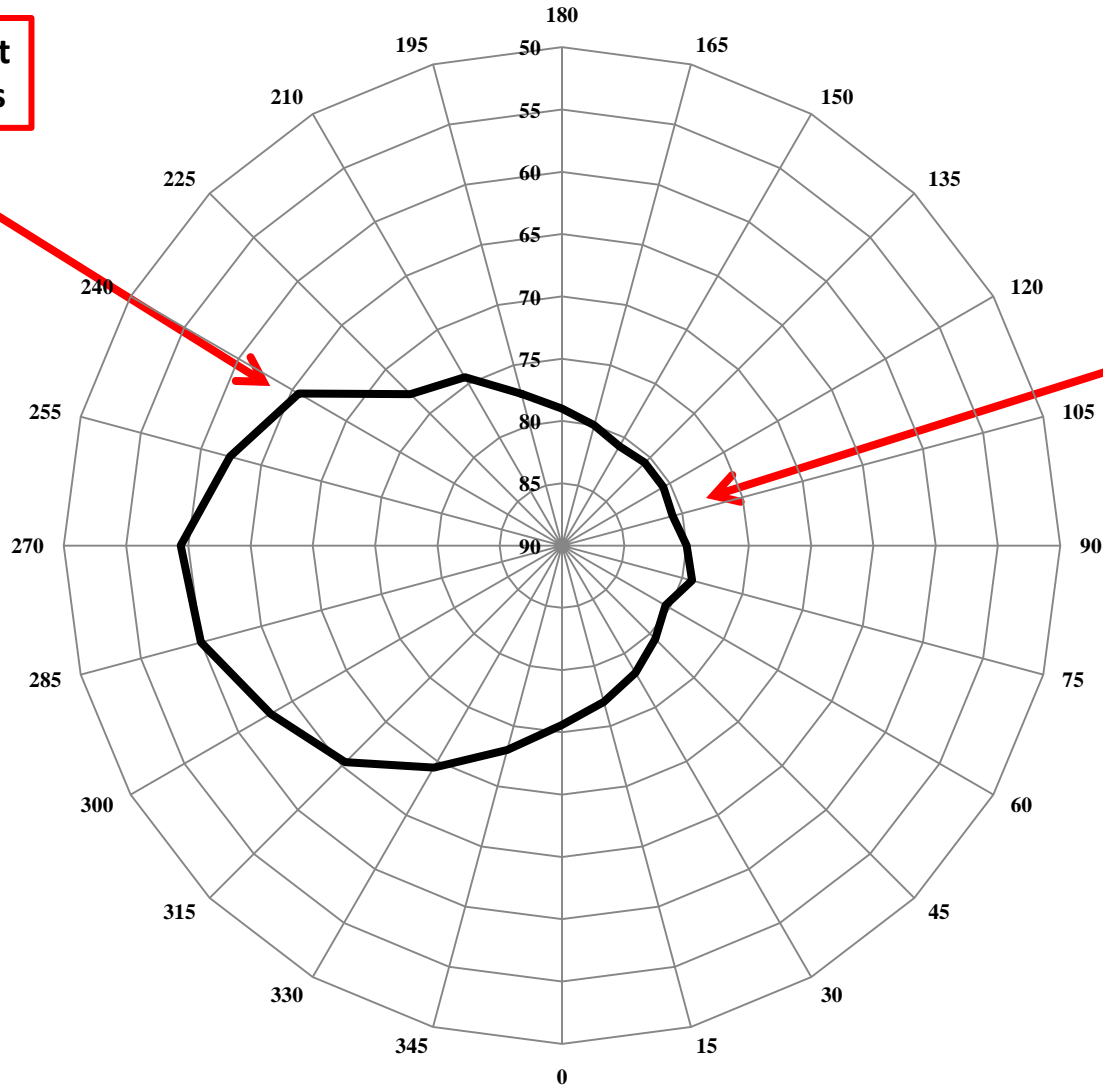
Example of Allowed Trajectory

Equatorward Extent of Polar Cap Absorption

Cutoff Latitude @ 0500 UT Under Quiet Conditions – Geographic Coordinates

Equatorward extent
of 1.0 MeV protons

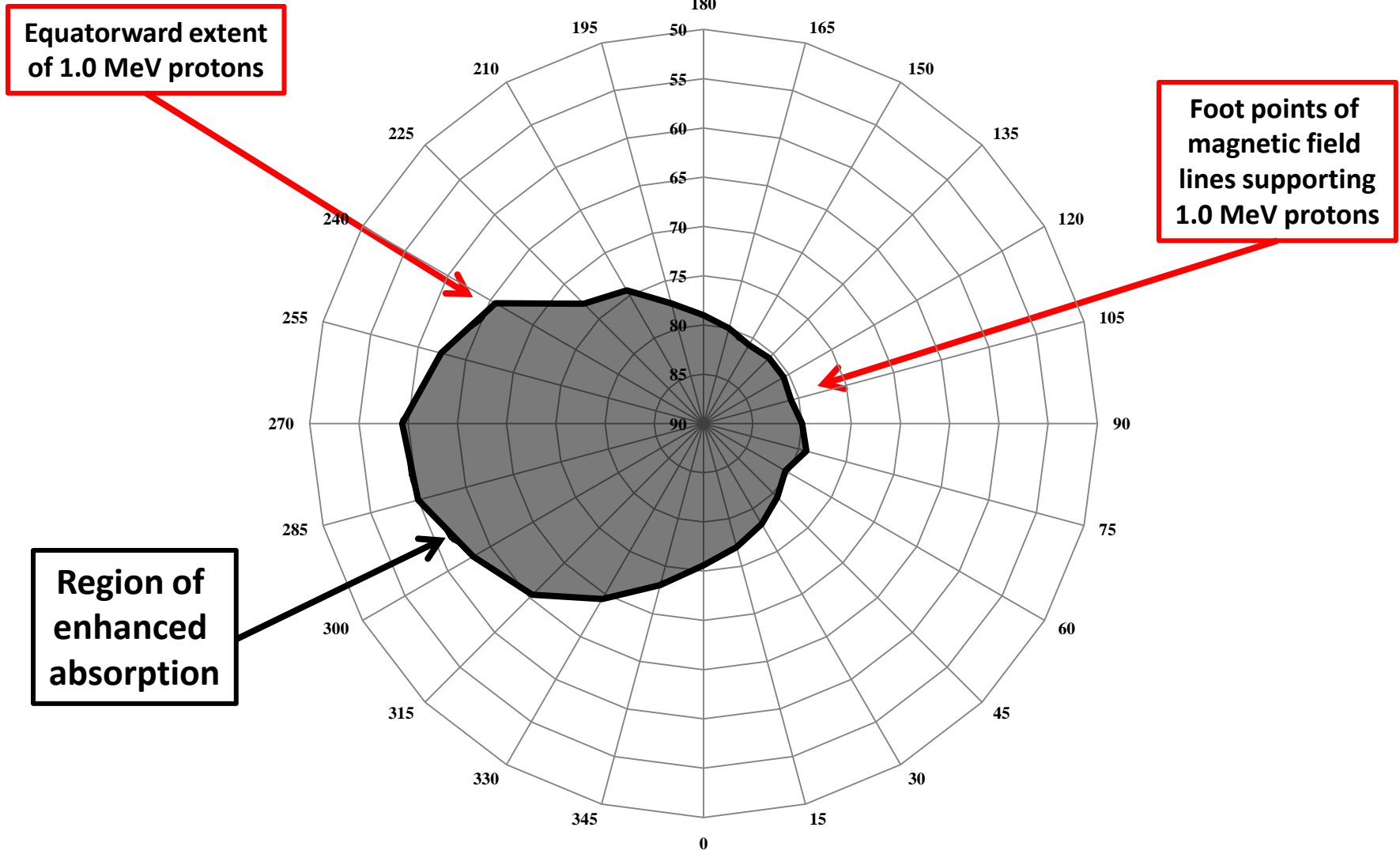
Foot points of
magnetic field
lines supporting
1.0 MeV protons



Cutoff latitude variation due to proton energy and Local Time

Equatorward Extent of Polar Cap Absorption

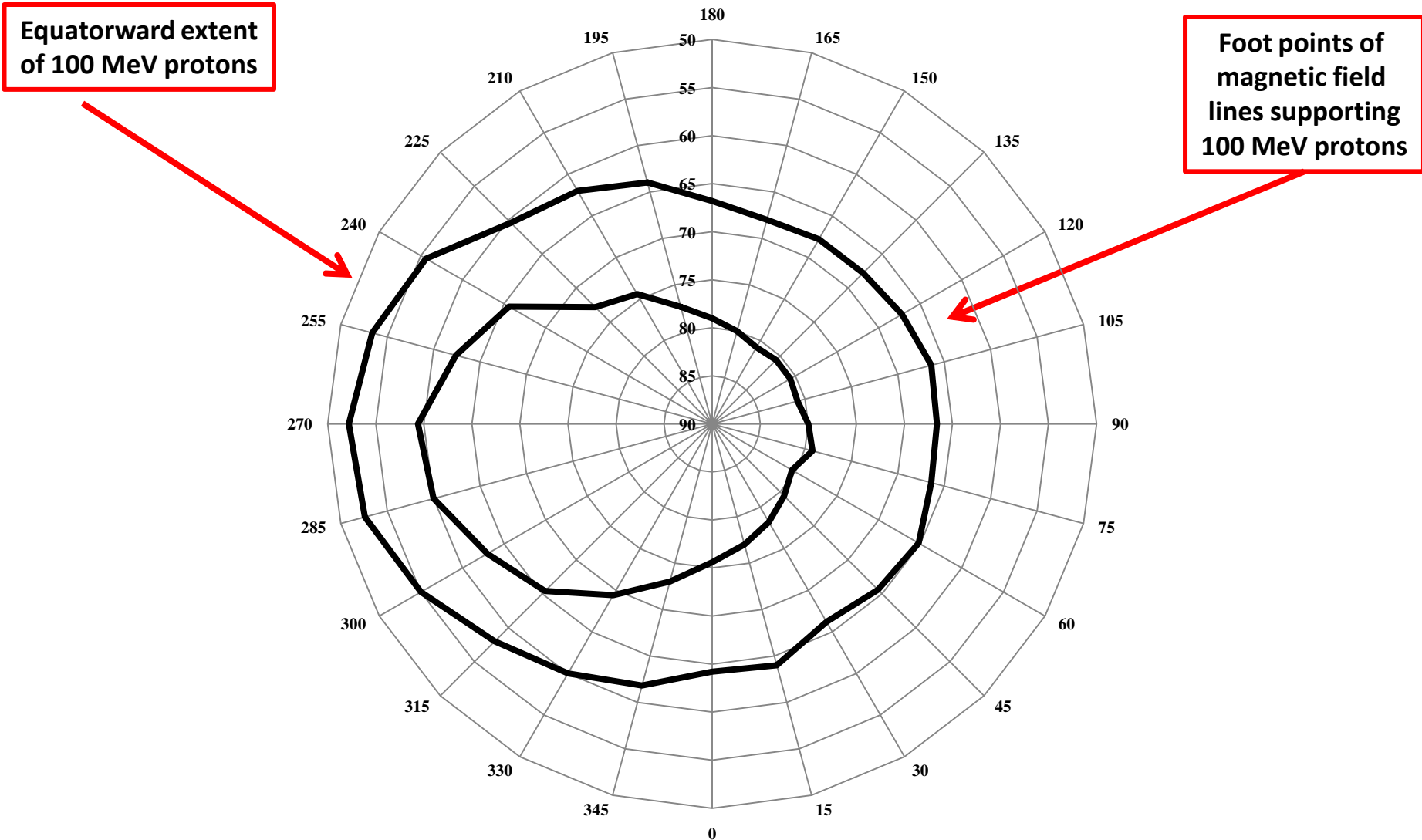
Cutoff Latitude @ 0500 UT Under Quiet Conditions – Geographic Coordinates



Cutoff latitude variation due to proton energy and Local Time

Proton Energy Cutoff Latitude

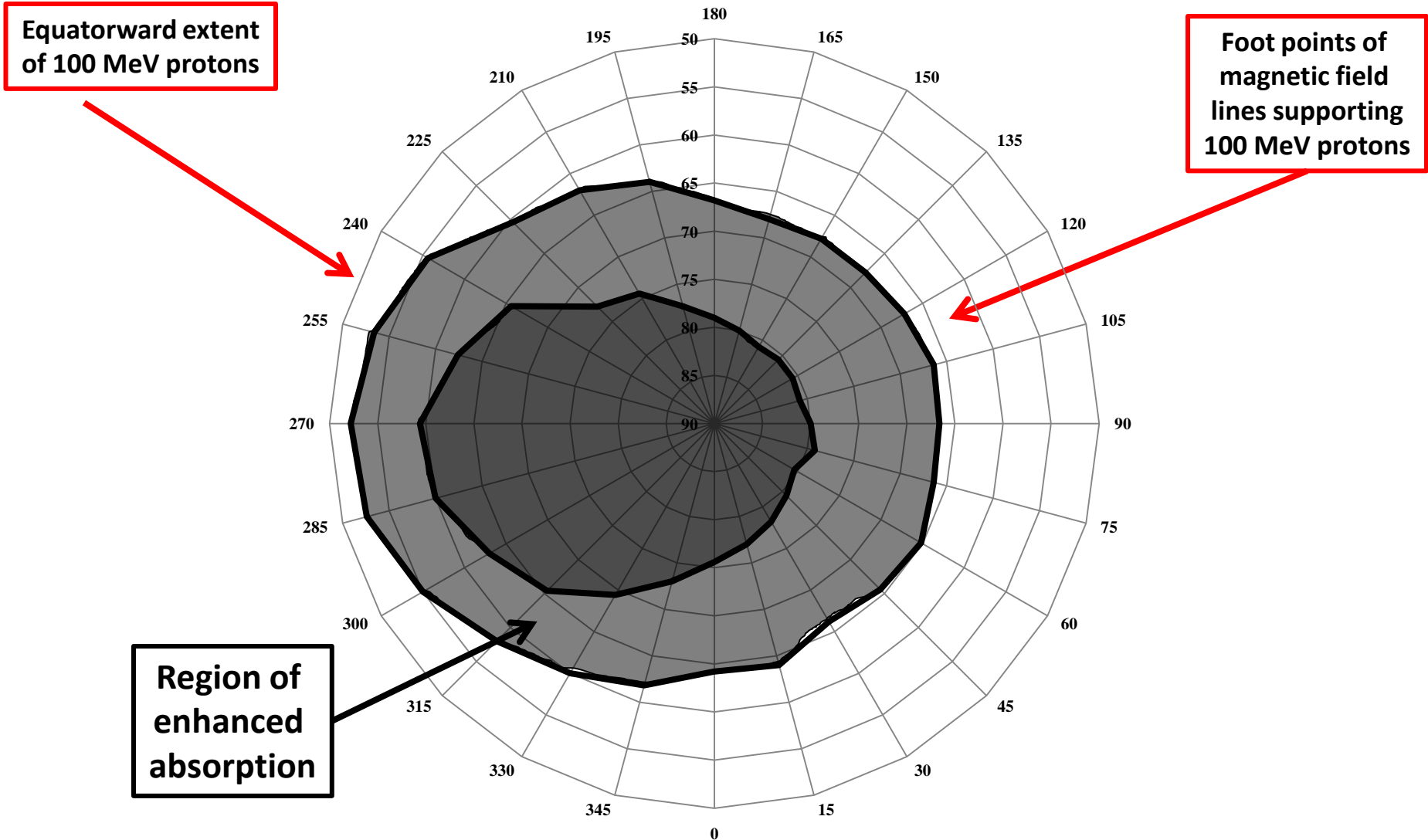
Cutoff Latitude @ 0500 UT Under Quiet Conditions – Geographic Coordinates



Cutoff latitude variation due to proton energy and Local Time

Proton Energy Cutoff Latitude

Cutoff Latitude @ 0500 UT Under Quiet Conditions – Geographic Coordinates



Cutoff latitude variation due to proton energy and Local Time

Proton Energy Cutoff Latitude

The Penumbra

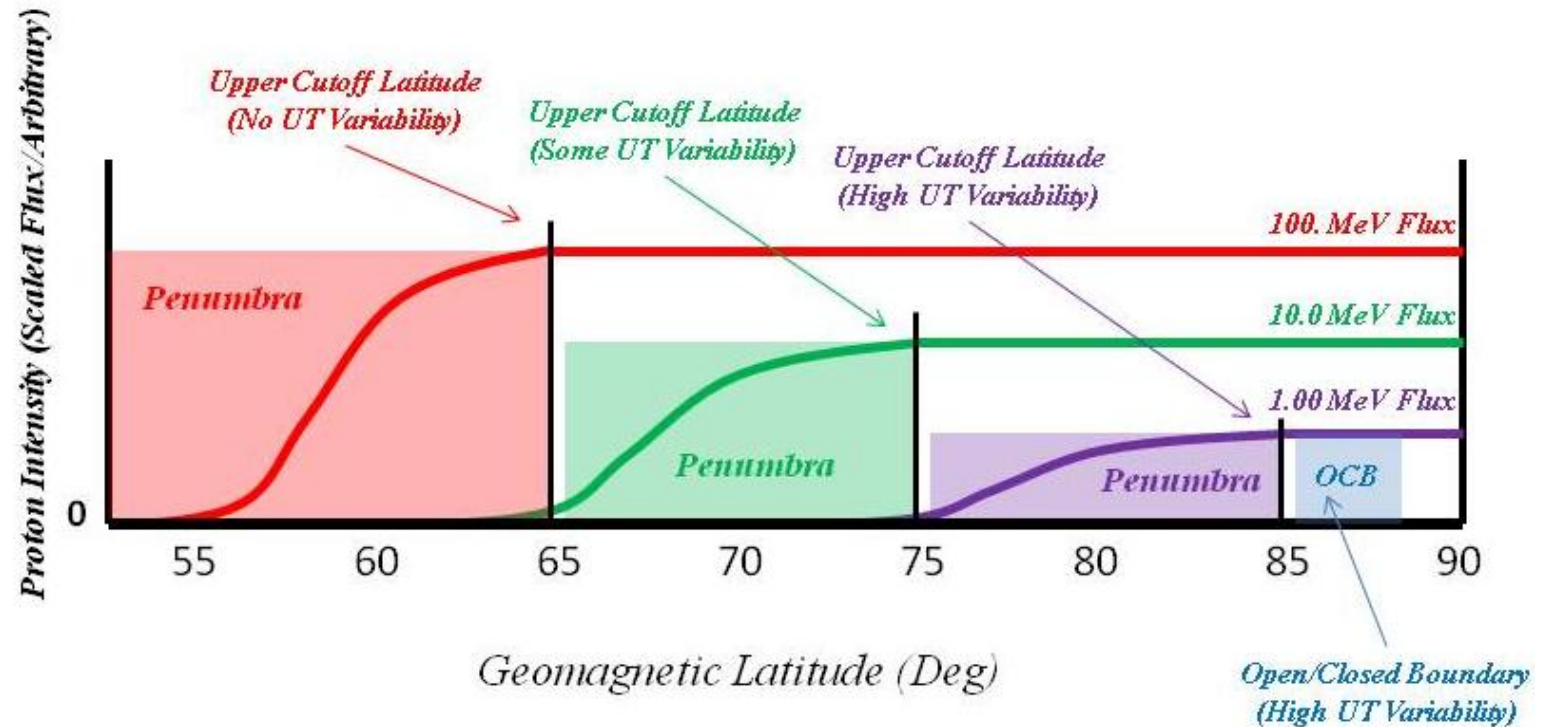
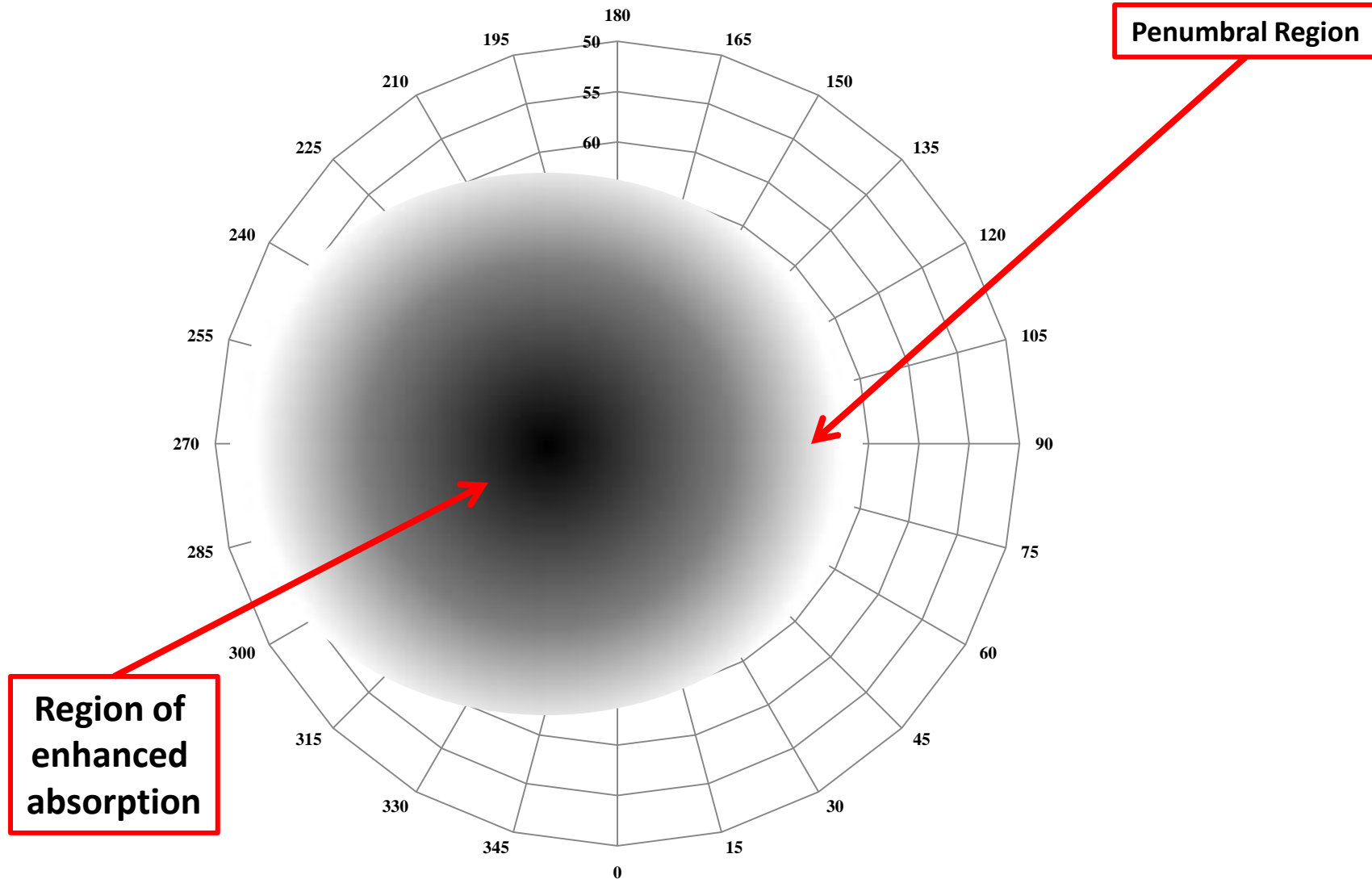


Illustration Only Not Drawn to Scale

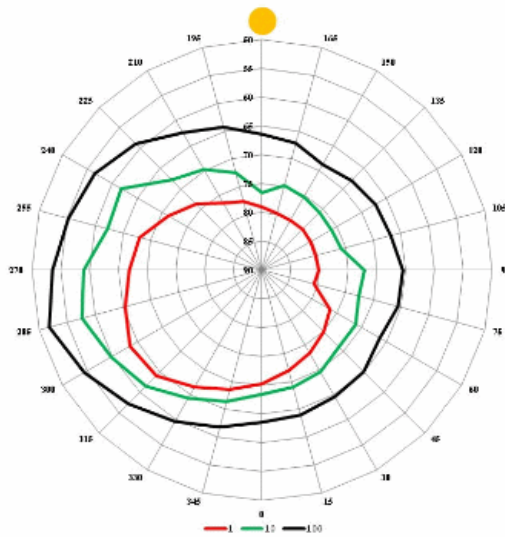
Proton Energy Cutoff Latitude

Cutoff Latitude @ 0500 UT Under Quiet Conditions – Geographic Coordinates

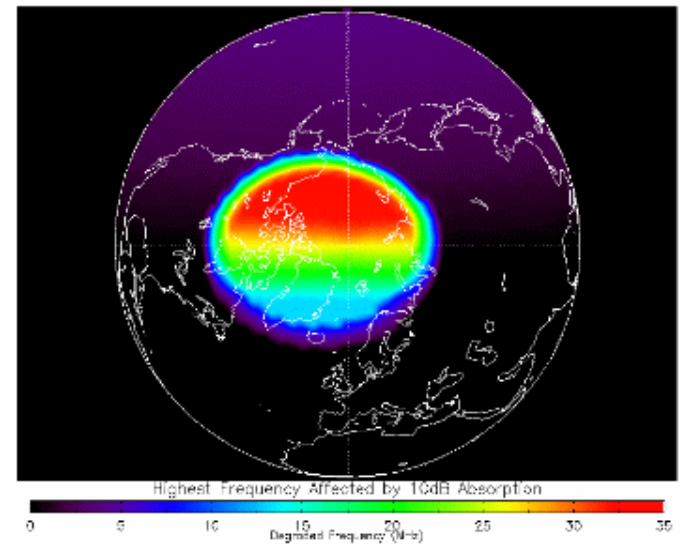


Cutoff latitude variation showing penumbral effect

Proton Energy Cutoff Latitude



imgflip.com

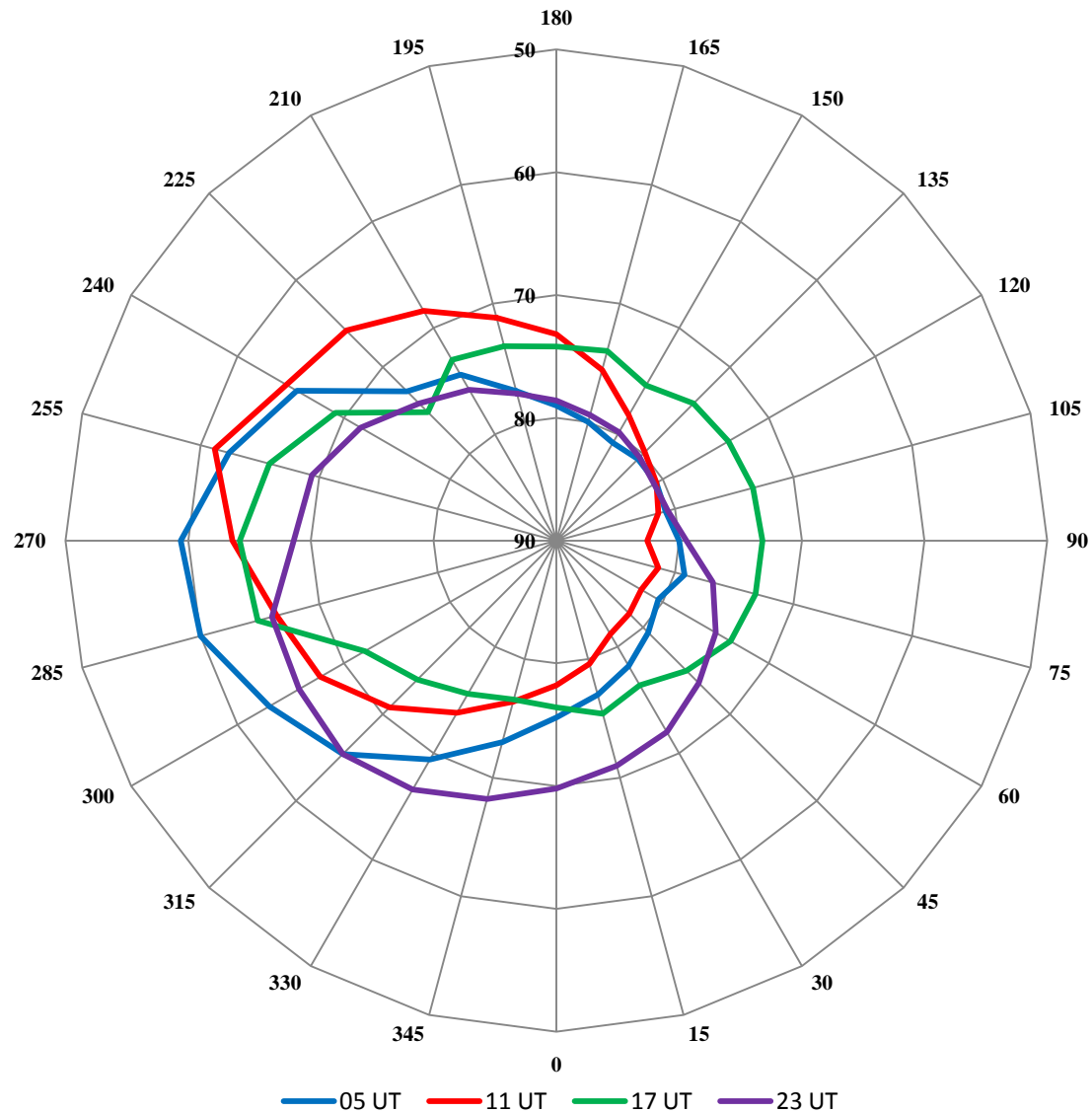


Excluded X-ray flux
Product Valid At : 2017-09-11 00:00 UTC
Moderate Proton Flux
NOAA/SWPC Boulder, CO USA

imgflip.com

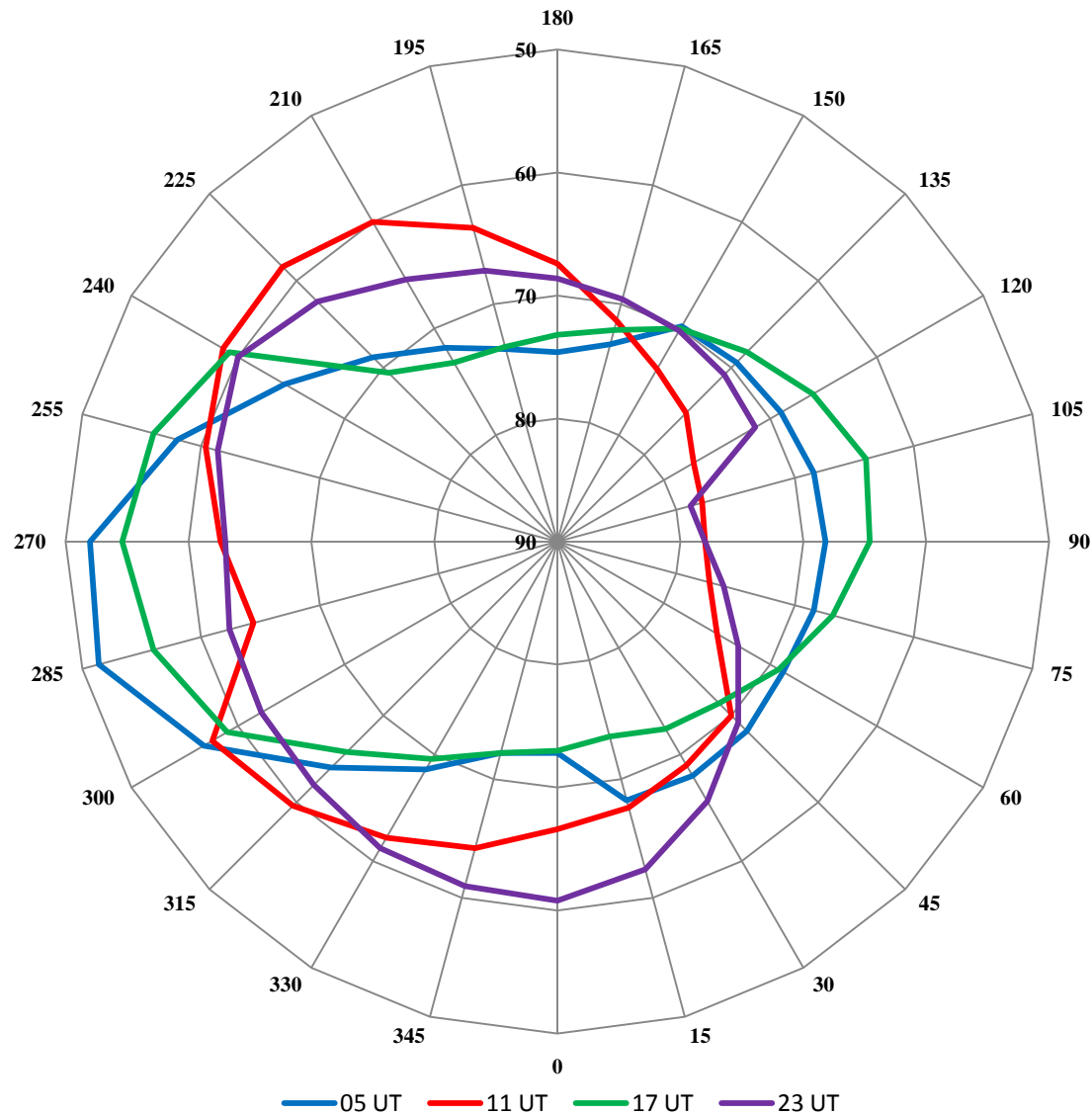
Proton Energy Cutoff Latitude

Quiet Conditions – 1 MeV Proton

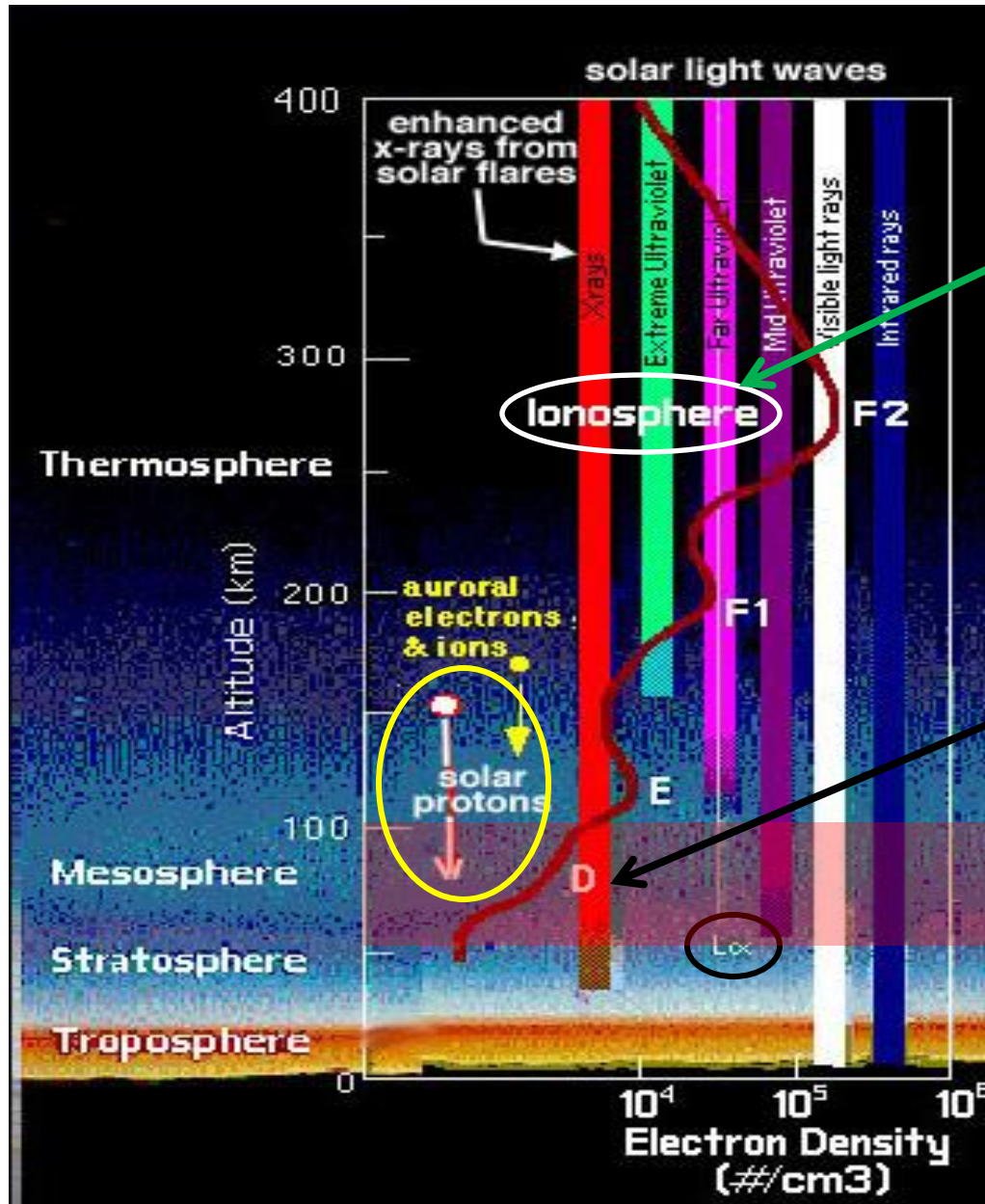


Proton Energy Cutoff Latitude

Severe Conditions – 1 MeV Proton



HF Absorption – Normal Conditions



Photons collide with neutral atoms and molecules creating ionosphere (free electrons)

Photons penetrate to different depths

D-Region

$$\text{Absorption} \propto \frac{N \cdot \nu}{f^2}$$

$N \equiv$ Electron density

$\nu \equiv$ Collision frequency

$f \equiv$ Radio wave frequency

Why Is This Important...

High Frequency radio wave propagation is important...

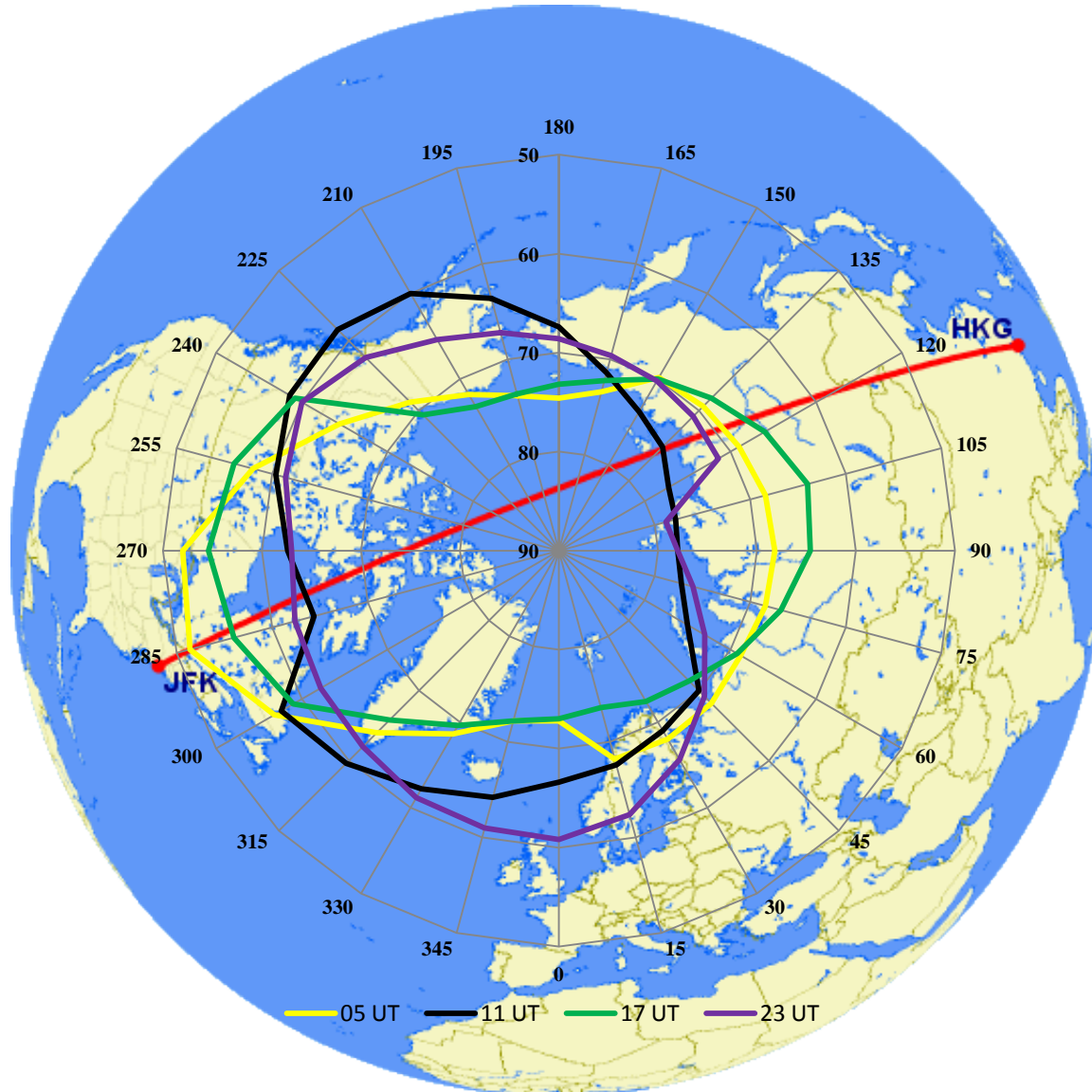
- Commercial and Military Aviators
 - Over 7000 transpolar commercial flight per year
- RADAR Installations
- Anyone using transpolar HF radio wave propagation

Why Is This Important...



1 MeV Proton Cutoff Boundary, Quiet Conditions

Why Is This Important...



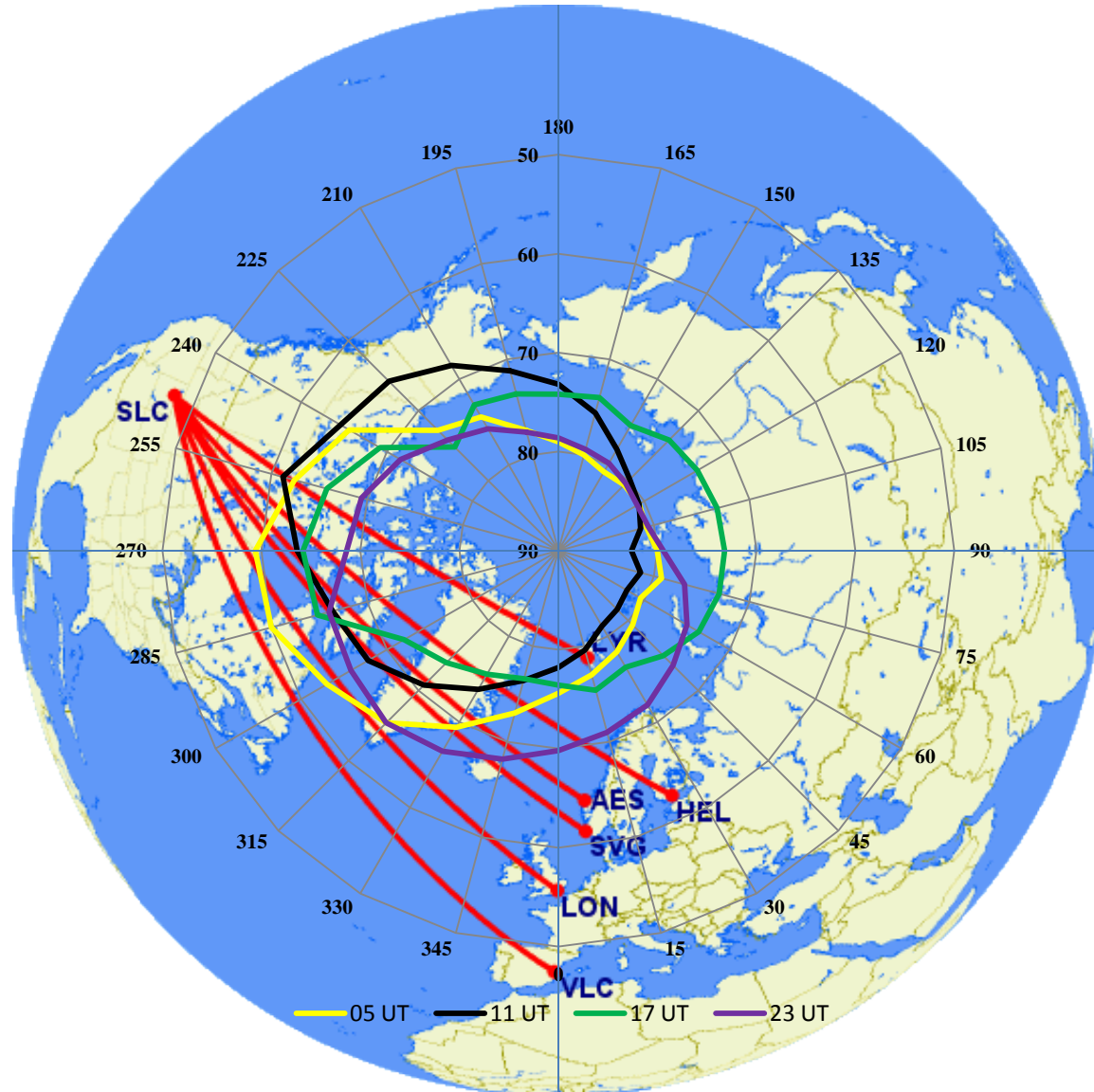
1 MeV Proton Cutoff Boundary, Severe Conditions

Amateur Radio Involvement...

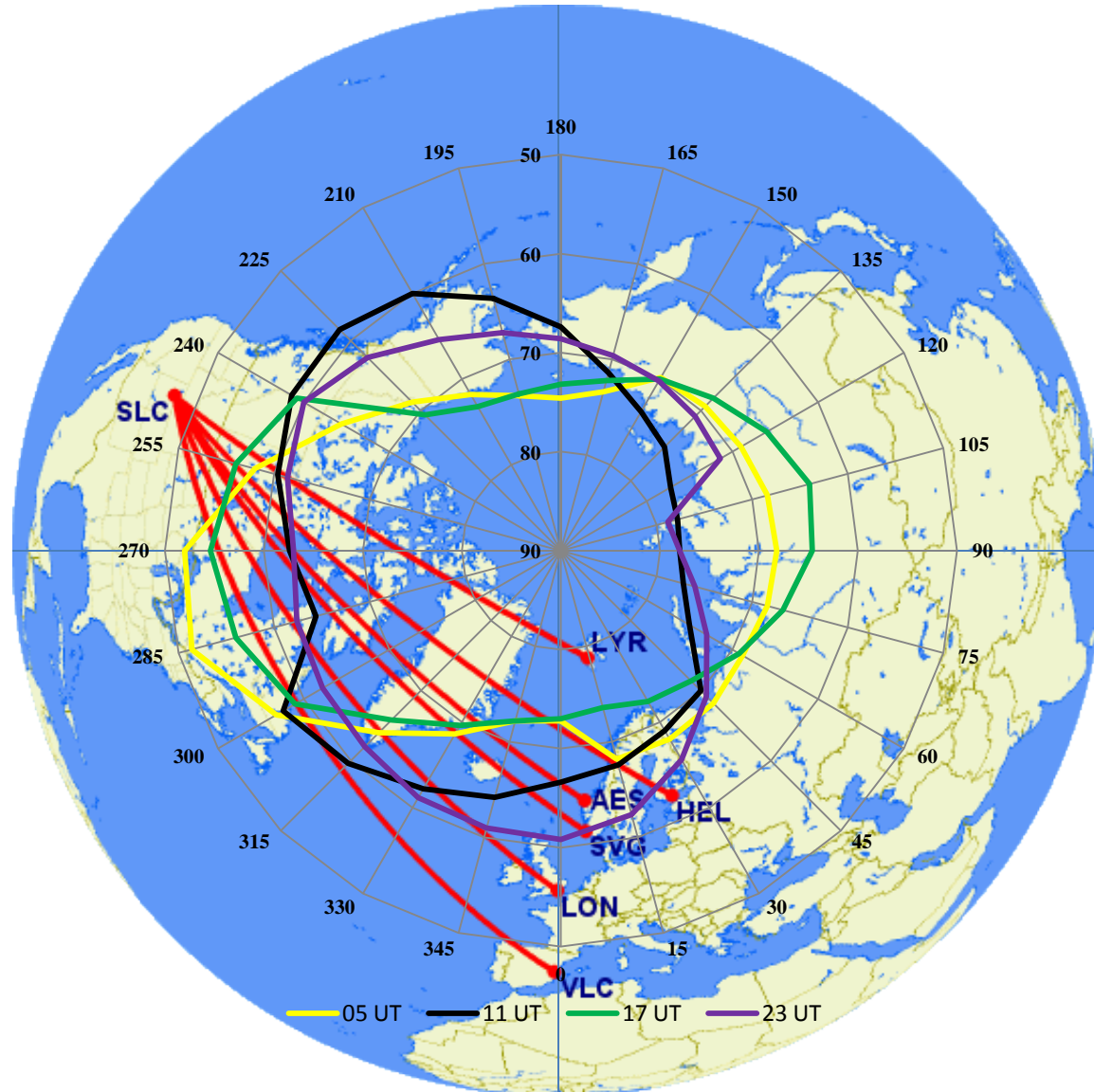


1 MeV Proton Cutoff Boundary, Quiet Conditions

Amateur Radio Involvement...



Amateur Radio Involvement...



1 MeV Proton Cutoff Boundary, Severe Conditions

Summary

- **Modeled Cutoff Latitude Shows Diurnal Variation
aka “UT Effect”**
- **UT Effect an important consideration:**
 - **Transiting Polar Regions**
 - **Modeling D-region characteristics**
 - **HF absorption**
- **Amateur Radio Could Help Demonstrate UT Effect**
- **Enhanced Understanding/Improvement of PCA Effects**

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<https://ccmc.gsfc.nasa.gov/models/modelinfo.php?model=Tsyganenko%20Magnetic%20Field>